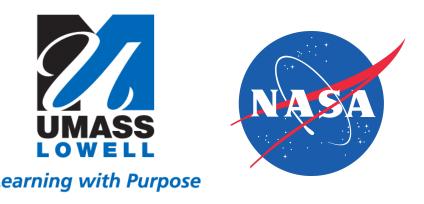


Microwave Kinetic Inductance Detectors for the PICTURE-C Mission



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The Planetary Imaging Concept Testbed Using a Recoverable Experiment – Coronagraph (PICTURE-C) will directly image debris disks and exozodiacal dust around nearby stars from a high-altitude balloon using a vector vortex coronagraph and Microwave Kinetic Inductance Detectors (MKIDs). We present the design and preliminary characterization of the MKID instrument. The cryostat uses liquid helium and an Adiabatic Demagnetization Refrigerator to maintain a detector temperature of 100 mK, well below the superconducting transition of the Platinum Silicide MKIDs. The 10,000 pixel MKID array will be optimized for 600nm with a 20% bandpass, and will be read out by ten ROACH2 and custom ADC/DAC boards.

Introduction

Planetary Imaging Concept Testbed Using a Recoverable Experiment – Coronagraph (PICTURE-C) is a mission that will use adaptive optics, a coronagraph, and an optical MKID camera to observe and characterize the scattered light from nearby debris disks

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Superconductor	Platinum Silicide	
Pixel Count	10,000	
Operating Temperature	100 mK	
Maximum Count Rate	~2500 photons/sec (soft limit)	
Resonator Q _i	~100,000	
# of microwave feedlines	5	
Frequency Range	4-8 GHz	
Photon timing resolution	~1 μs	

Readout

- Pulse detection readout based on DARKNESS firmware
- 10 ROACH2 boards
- 10 custom ADC/DAC boards

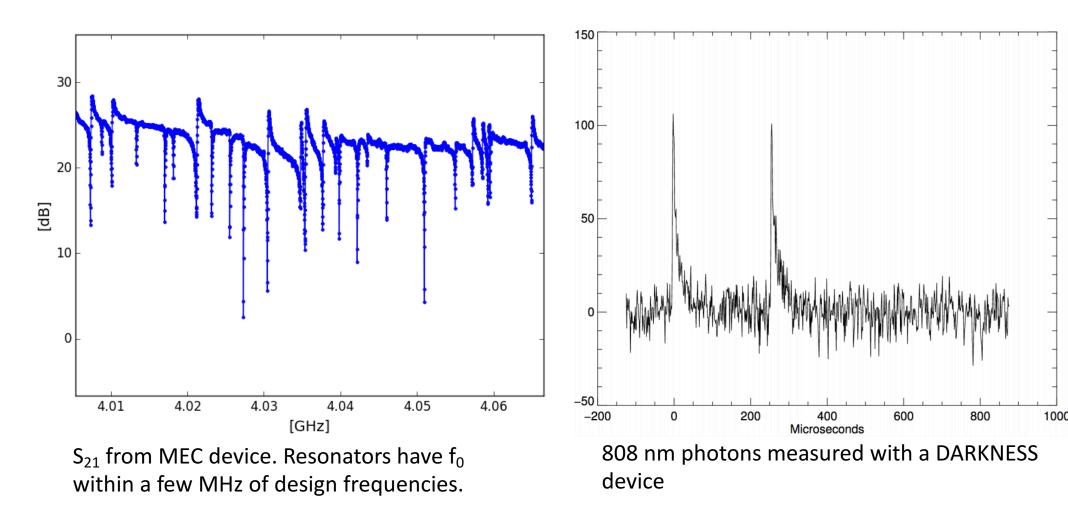
UCSB is constructing the MKID camera

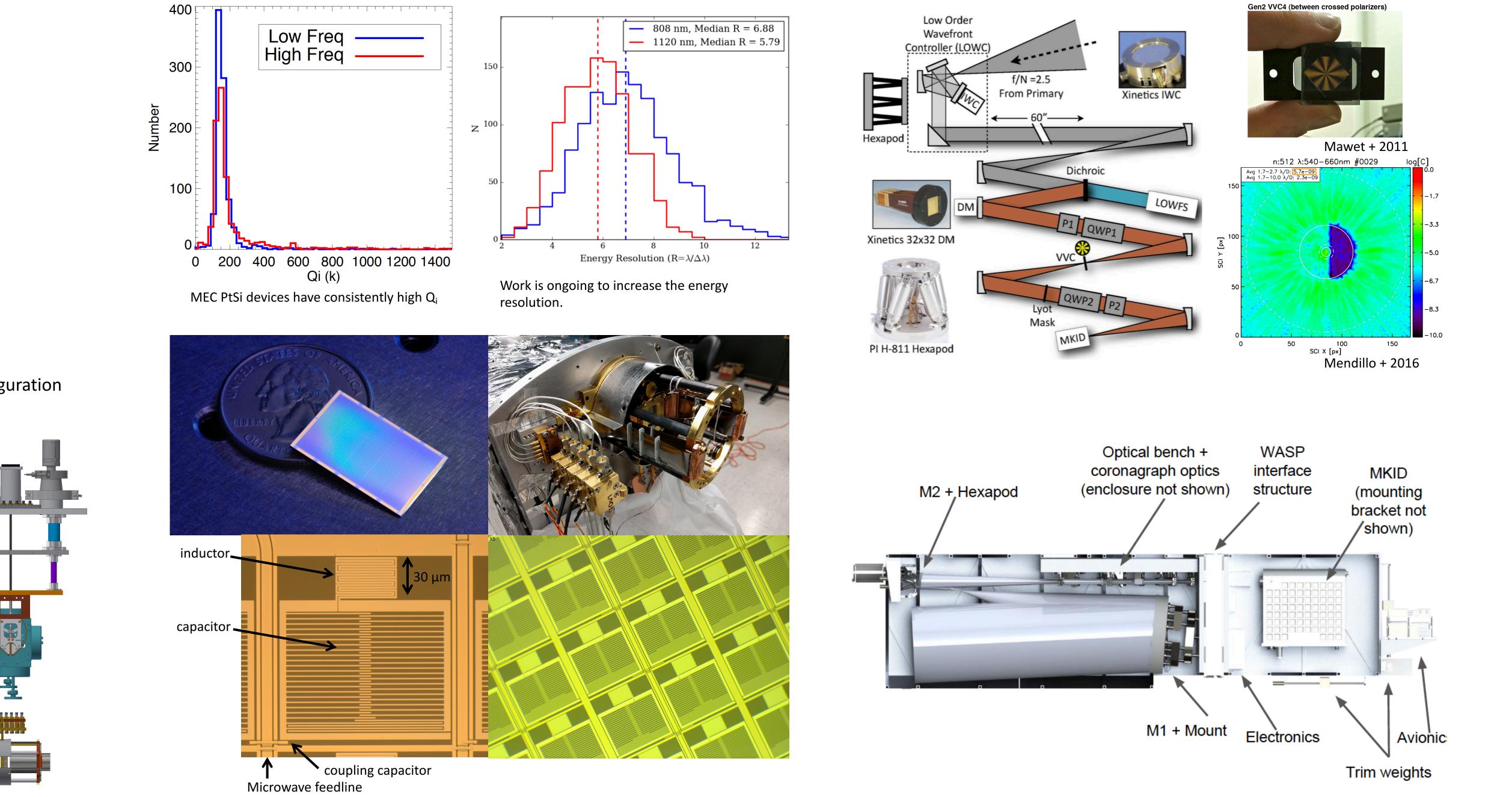
Wavelength	600 nm
Bandpass	20% (540nm to 660nm)
Contrast at IWA	10-7
Resolution ($\lambda/\Delta\lambda$)	20
Control Region	Half-plane from 1.7 to 10 λ/D (0.35" to 2.1")

Cryostat

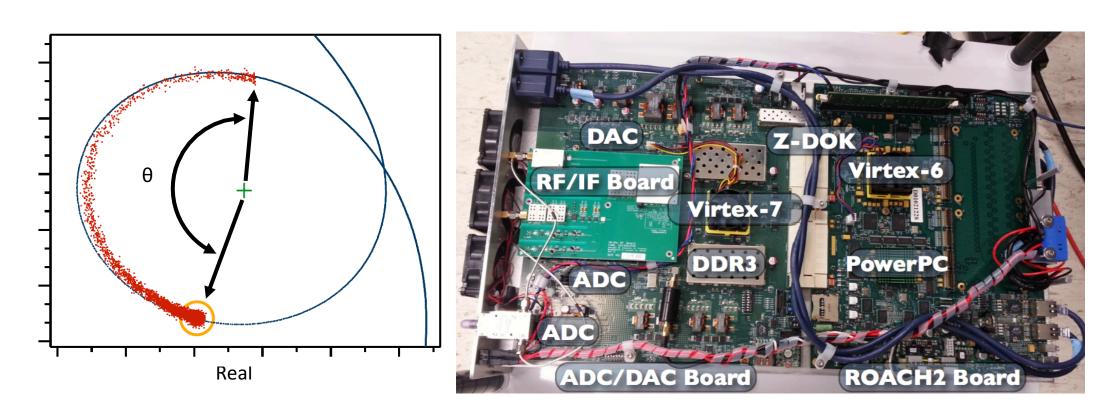
- Liquid helium/nitrogen
- Designed hold time ~50 hrs
- Adiabatic Demagnetization Refrigerator
- ~115 kg full
- NbTi flex cables from 4 K to 100 mK for low loss and low thermal loading
- 5 HEMTs with ~40 dB of gain at 4 GHz
- Pulse tube cooler for in-lab testing







- 10 custom IF boards
- ~1400 W, will require large radiator panel



Telescope and Optics

- 60 cm off-axis paraboloid primary
- 2 deformable mirrors will actively maintain a "dark hole" in the focal plane
- Vector vortex coronagraph will suppress on-axis starlight

