

Neutrino masses

In SM: $\lambda_{ij}^e \phi l_i \bar{e}_j$

↑

diagonalize
using $U_L^l \times U_R^e$

No $\bar{\nu} \Rightarrow m_{\nu_i} = 0.$

$m_{\nu_i} \neq 0$: Experimental evidence for beyond SM

Recall SM has three "accidental" symms:

$l_i \rightarrow e^{i\alpha_i} l_i$ e, μ, τ lepton #.

$\bar{e}_i \rightarrow e^{-i\alpha_i} \bar{e}_i$ (anomalies violate \rightarrow
non pert. ...)

... interesting role...

Eg. if violate, mass w/out $\bar{\nu}$!

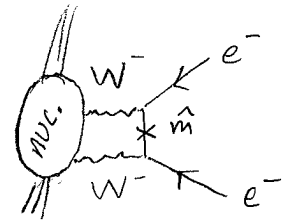
(unique feature - ν 's chargeless)

$$\mathcal{L}_{\hat{m}} = - \frac{\hat{m}_{ij}}{2} \nu_i \nu_j \quad (+ h.c.)$$

← understood
henceforth

recall violates $U(1)_{L_i}$

"Majorana mass." $\sim (Ov2\beta)$



4.2

Also breaks $SU(2) \times U(1)$ - need source

eg. in SM, $\tilde{\Phi} \cdot l_i$: singlet, $Y = 0$

$$\Rightarrow \mathcal{L}_{\tilde{\Phi} l^2} = - \frac{\hat{\lambda}_{ij}}{M} \tilde{\Phi} l_i \tilde{\Phi} l_j$$

... nonrenormalizable

or: can get from higher rep of $SU(2)$ - 3!

either way ... new physics

The old-fashioned way? $\bar{\nu}_i$ $Y = 0$

$$m_{ij}^{\nu} l_i \bar{\nu}_j$$

eg. SM

$$\lambda_{ij}^{\nu} \tilde{\Phi} l_i \bar{\nu}_j, \quad m_{ij}^{\nu} = \lambda_{ij}^{\nu} \frac{v}{\sqrt{2}}$$

$$\leadsto U(1)_{L_i} \text{ conserved, } \bar{\nu}_i \rightarrow e^{-ix_i} \bar{\nu}_i$$

But, data: m_ν 's $\lesssim .1 \text{ eV}$
 $\Leftrightarrow \lambda_{ij}^\nu \lesssim 10^{-13} - 10^{-12}$

... begs explanation.

new symmetries involving $\bar{\nu}$?

$\bar{\nu}$ is different - pure singlet

This \leadsto another new possibility:

$$\mathcal{L}_M = - \frac{M_{ij}}{2} \bar{\nu}_i \bar{\nu}_j \quad \text{gauge allowed}$$

(but: spoils $U(1)_{L_i}$'s)

Most general possibility:

$$\begin{pmatrix} \nu \\ \bar{\nu} \end{pmatrix} \cdot \underbrace{\begin{pmatrix} \hat{m} & m^T \\ m & M \end{pmatrix}}_M \begin{pmatrix} \nu \\ \bar{\nu} \end{pmatrix}$$

... and provides a candidate explanation for small masses!

Suppose $\hat{m} = 0$.

Mass e'states: diagonalize M .

M_{ij} in general is renormalized, might expect

$$M_{ij} \sim \Lambda_{NP}. \quad \text{If } M \gg m,$$

$$m_{\text{light}} \simeq -m M^{-1} m^T = \underline{\text{small}}$$

"Seesaw mechanism"

Effective field theory description:

$$\mathcal{L}_M = m_{ij} \nu_i \bar{\nu}_j + \frac{M_{ij} \bar{\nu}_i \bar{\nu}_j}{2} \quad \leftarrow \text{symmetric}$$

$\bar{\nu}$ heavy \rightsquigarrow integrate out: $\delta/\delta\bar{\nu}_j$

$$0 = \nu^T M + \bar{\nu}^T M \quad (\text{matrix notation})$$

$$\text{or } \bar{\nu}^T = -\nu^T m M^{-1}$$

$$\rightsquigarrow \mathcal{L}_{\text{eff}} = -\frac{\nu^T m M^{-1} m^T \nu}{2} \quad \checkmark$$

So: Maj. mass for ν (or $\frac{(\tilde{\phi} l)^2}{M}$)

Effective description of theory w/ heavy $\bar{\nu}$.

(either way, violate $U(1)_{L_i}$)

If so, what's M ?

$m \sim \lambda v$; suppose $\lambda \sim \mathcal{O}(1)$
(compare, e.g., top quark)

$$m_\nu \lesssim .1 \text{ eV} \Rightarrow M \sim \frac{v^2}{1 \text{ eV}} \sim \text{few} \times 10^{14} \text{ GeV.}$$

a new heavy scale. just below M_{GUT} (... soon),
(note smaller λ doesn't get closer to M_{GUT})

Mixing (the evidence for m_ν 's) :

Weak current:
$$J^\mu = \begin{pmatrix} \nu_i \\ e_i \end{pmatrix}^\dagger \frac{\tau^a}{2} \bar{\sigma}^\mu \begin{pmatrix} \nu_i \\ e_i \end{pmatrix}$$

~ CKM story:

1) Diagonalize m^e : U_L^e, U_R^e

2) Then have $-\hat{M}_{ij} \nu_i \nu_j$

$$\hat{M}_{ij} = \begin{matrix} \text{symmetric} \\ \text{complex} \end{matrix} = U^\dagger m_i U \begin{matrix} \uparrow \\ \text{diag} \end{matrix}$$

U_{PMNS} Pontecorvo - Maki - Nakagawa - Sakata.

$$\begin{matrix} \text{weak } e' \text{ states} \rightarrow & \nu' = U_{\text{PMNS}} \nu & \leftarrow \text{mass } e' \text{ states.} \end{matrix}$$

CKM : 3 angles, 1 phase

(ex)

PMNS : 3 angles, 3 phases

E.g. parametrize as

$$U_{\text{PMNS}} = U_{23}(\theta_{23}) U_{13}(\theta_{13}, \delta) U_{12}(\theta_{12}) \times \begin{pmatrix} 1 & 0 & 0 \\ 0 & e^{i\phi_1} & 0 \\ 0 & 0 & e^{i\phi_2} \end{pmatrix}$$

\uparrow rot in 2-3 plane \uparrow Dirac ~~CP~~ phase \uparrow Majorana ~~CP~~ phases

Experiments

1) Sun : makes ν_e 's.

we don't see enough. big puzzle for a long time.
 (ν flux detection: Nobel '02, Davis/Koshiba)

$$\theta_{12} + \text{density effects} \rightsquigarrow \nu_e \rightarrow \nu_\mu$$

\uparrow
 can now see.

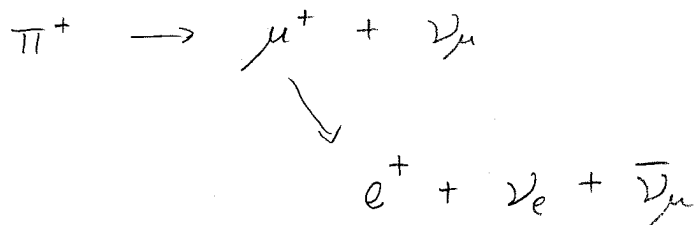
also reactor ν 's : KamLAND

$$\rightsquigarrow \theta_{12} \approx 33.9 \pm 1.6^\circ \quad (\text{big!})$$

also

$$\Delta m_{12}^2 \approx (7.9 \pm .4) \times 10^{-5} \text{ eV}^2$$

2) Atmospheric cosmic p's + atmosphere \rightarrow π 's.



π^- : charge conjug.

$$\Rightarrow R = \frac{N_{\nu_\mu}}{N_{\nu_e}} \sim 2. \quad R_{obs} \sim .6$$

interp: $\nu_\mu \rightarrow \nu_\tau$

+ K2K, MINOS, ... $\sin^2 \theta_{23} \approx .47$

i.e. $\theta_{23} \approx \frac{\pi}{4}$ - maximal!

$$|\Delta M_{32}^2| \approx (2.4 \pm .3) \times 10^{-3} \text{ eV}^2$$

3) CHOOZ: $\sin^2 \theta_{13} \lesssim 0.05$

etc.

* All well-established data can be explained by mixing via UPMNS

* Might consider other alternatives

$\Gamma_Z \sim N_\nu = 3$; but sterile (singlet)
can \leadsto large mixing

* Whatever new physics is, it ultimately needs to explain ν -masses.