HST, super-BMS, Feynman Diagrams and Black Holes

Tom Banks (work with W.Fischler)

KITP Quantum Gravity, May 18, 2014

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 Multiple Quantum Systems Related by Equivalence of Density Matrices on Overlaps.

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- Evolution Operator Maps Past BMS Onto Future BMS.

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The Super BMS Algebra

$$[Q_{\alpha}(P,m), \bar{Q}_{\beta}(Q,n)]_{+} = \gamma^{\mu}_{\alpha\beta}P_{\mu}\delta(P \cdot Q)Z_{mn}. P^{2} = Q^{2} = 0,$$

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unless f vanishes outside finite number of spherical caps for P > 0, and in annuli around those caps for P = 0. Two different algebras for $\frac{P_0}{|P_0|} = \pm 1$.

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• S-matrix maps them into each other $SQ^- = Q^+S$.

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- In d dimensions, holoscreen spinor bundle with eigenvalue cutoff n : ψ_a^(k₁...k_{d-2}).
- ► $M_k^m \equiv \psi_{kl}^a {}^{\dagger} \psi_a^{ml}$, *I* a *d* 3 anti-symmetric tensor index. When acting on scattering states: for $n = N \rightarrow \infty$, *M* has blocks of size $1 \ll K_i \ll N$, plus one large block of size $N \sum K_i$.

▶ For fairly general Hamiltonians of the form $H_{in}(n) = \sum_{i} P_0^i + \frac{1}{n^{2(d-3)}} \text{Tr } P(M)$, with *P* a polynomial of degree n^{d-4} .

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• $P_0 = \sum P_0^i \propto \sum K_i^{d-3}$ asymptotically conserved.

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- Large distance eikonal scattering has scaling with energy and impact parameter of "Newton's Law".
- Many amplitudes can be represented as space-time diagrams with localized vertices, as a consequence of consistency conditions for distant trajectories.

Black Holes

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- Probability of emitting jets from such states is thermal with $T \sim n^{-1}$, so they behave like black holes.

 All of these models behave qualitatively like a model of QG, but most are not Lorentz invariant as K_i go to infinity.
 Enough parameters, plausibly, to tune for Lorentz invariance.

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 QUEFT is a good approximation only in regimes where particles are decoupled from the horizon.

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Please Join In.