

THE SHAPE OF ORBITS IN THE SCHWARZSCHILD GEOMETRY

This is a *Mathematica* program to compute and display the shapes of orbits in a Schwarzschild geometry. The Schwarzschild radial coordinate is measured in units of M , so that $M=1$ in the following formulae. Throughout the variable $u=r/M$ is used. Here is a list of what you must set to run the program:

- (1) The angular momentum: ℓ .
- (2) The energy parameter: $\mathcal{E} = (e^2 - 1)/2$.
- (3) The starting radius for an orbit which is not bound: rst .
- (4) The number of orbits to be computed if the orbit is bound: $norbit$.

These are set by editing the definition statements at various places in the program. **You have to make sure these parameters are set so the orbit is classically allowed and doesn't start at a position where the value of the effective potential that is greater than \mathcal{E} .**

■ The potential:

The effective potential for radial motion V_{eff} given in (9.28) is here denoted simply by V . To be slightly more general a parameter **signewt** is introduced which multiplies the non-Newtonian $1/r^3$ term in the potential. Set it equal to 1 for general relativity, 0 for a Newtonian $1/r$ potential, and an appropriate value for a Newtonian $1/r$ potential with additional quadrupole moment term.

```
In[47]:= signewt = 1.  
Out[47]= 1.  
  
In[48]:= V[u_, \ell_] := -u + \ell^2 u^2 / 2 - signewt \ell^2 u^3
```

■ Specifying the orbit:

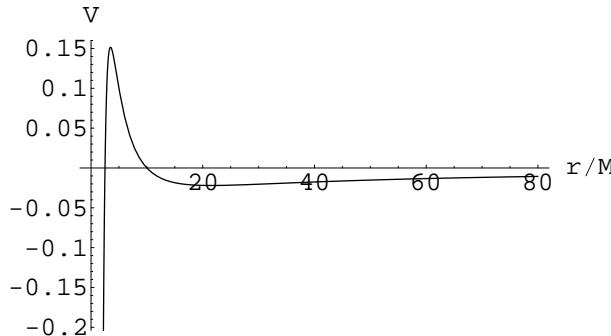
It takes four numbers to specify an orbit: (1) the angular momentum ℓ , (2) the energy parameter \mathcal{E} , (3) the starting radius rst , and (4) the number of orbits to be calculated, $norbit$.

(1) Specify the **angular momentum ℓ** :

```
In[49]:= \ell = 5  
Out[49]= 5
```

This is a graph of the effective potential vs r/M for this value of ℓ :

```
In[50]:= veff = Plot[V[1/r, t], {r, 2, 80}, AxesLabel -> {"r/M", "V"}]
```



```
Out[50]= - Graphics -
```

Next the values of the potential at its extrema are computed.

```
In[51]:= dv[u_, t_] := -1 + t^2 u - 3 signewt t^2 u^2
```

```
maxmin = NSolve[dv[ex, t] == 0, ex]
```

```
Out[52]= {{ex -> 0.0464816}, {ex -> 0.286852}}
```

```
In[53]:= vmin = V[ex /. maxmin[[1]], t]
```

```
Out[53]= -0.0219855
```

```
In[54]:= vmax = V[ex /. maxmin[[2]], t]
```

```
Out[54]= 0.151615
```

(2) Specify the **energy parameter** \mathcal{E} . It must be greater than the minimum value of the potential given above.

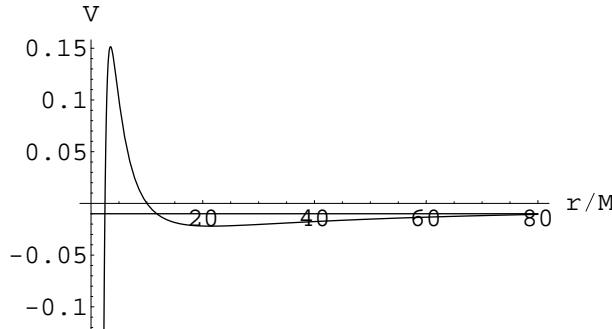
```
In[55]:= E = -.01
```

```
Out[55]= -0.01
```

Next we plot the effective potential and the energy parameter \mathcal{E} on the same graph:

```
In[56]:= sce := Plot[{E, veff}, {r, 0, 80}, DisplayFunction -> Identity]
```

```
In[57]:= Show[veff, sce, DisplayFunction -> $DisplayFunction]
```



```
Out[57]= - Graphics -
```

From this information you can see what kind of orbit it will be — bound orbit, plunge orbit, or scattering orbit.

Now the program evaluates the three turning points associated with this value of the energy parameter:

```
In[58]:= soln = NSolve[V[tp, t] == ε, tp]
```

```
Out[58]= {{tp → 0.0116597}, {tp → 0.0850696}, {tp → 0.403271}}
```

```
In[59]:= tp1 = tp /. soln[[1]]
```

```
Out[59]= 0.0116597
```

```
In[60]:= tp2 = tp /. soln[[2]]
```

```
Out[60]= 0.0850696
```

```
In[61]:= tp3 = tp /. soln[[3]]
```

```
Out[61]= 0.403271
```

(3) Specify the **starting radius** `rst`. This together with the energy parameter determines what kind of orbit is computed — bound, asymptotic from and to infinity, plunge, emerging from $r=2M$, etc. depending on where it is set relative to the above turning points. However, only if the orbit goes off to infinity is the *value* of `rst` used.

```
In[62]:= rst = 20.
```

```
Out[62]= 20.
```

```
In[63]:= ust = 1 / rst
```

```
Out[63]= 0.05
```

(4) Finally specify the **number of orbits**, `norbit`, if the orbit is bound.

```
In[64]:= norbit = 3
```

```
Out[64]= 3
```

■ Computing the Orbit

We are now going to integrate $\{(\ell/2^{1/2})[E-V[u,\ell]]^{-1/2}\}$ to find the orbit in the form $\phi(r)$. The integrand diverges near the turning points, so we specify a parameter **eps** which determines how close the numerical integration comes to the turning points.

```
In[65]:= eps = .00000001
```

```
Out[65]= 1. × 10-8
```

Next the program picks the appropriate radii to start and end the integration of the orbit given the energy parameter and starting radius. There are four kinds of orbits.

- (a) Bound orbits which start at the outer turning point.
- (b) Orbits which come in from infinity and go out again. These start at r_{st} .
- (c) Orbits which start close to $r=2M$ and fall back to it.
- (d) Plunge orbits which start at r_{st} and end at $r=2M$. For bound the starting radii are the

Tests for bound orbits:

```
In[66]:= testa = (ε < 0) && (ust < tp2)
```

```
Out[66]= True
```

```
In[67]:= If[testa, u1 = tp1 (1 + eps); u2 = tp2 (1 - eps)]
```

```
Out[67]= 0.0850696
```

Tests for orbits coming from and going to infinity:

```
In[68]:= testb = (ε > 0) && (ε < vmax) && (ust < tp2)
```

```
Out[68]= False
```

```
In[69]:= If[testb, u1 = ust; u2 = tp2 (1 - eps)]
```

Tests for orbits which start from $r=2M$ and go back to it:

```
In[70]:= testc = (ε < vmax) && (ust > tp3)
```

```
Out[70]= False
```

```
In[71]:= If[testc, u1 = .5; u2 = tp3 (1 + eps)]
```

Tests for plunge orbits coming from infinity going to $r=2M$:

```
In[72]:= testd = (ε > vmax)
```

```
Out[72]= False
```

```
In[73]:= If[testd, u1 = ust; u2 = .5 (1 - eps)]
```

If the parameters are set correctly only one of the four options should be “true” and the other three should be “false”. The values of $u1$ and $u2$ are displayed below:

```
In[74]:= u1
Out[74]= 0.0116597

In[75]:= u2
Out[75]= 0.0850696
```

A number of functions are defined, the chief of which is the angle **theta** which is the angle ϕ swept out from the innermost turning point. This is given by the following numerical integral:

```
In[76]:= theta[u_, ε_, ℓ_, u1_] := NIntegrate[(ℓ/2^(1/2)) (ε - V[w, ℓ])^(-1/2), {w, u1, u}]
```

delphi is the total angle ϕ swept out between two endpoints of the orbit:

```
In[77]:= delphi = theta[u2, ε, ℓ, u1]
```

```
Out[77]= 3.73604
```

z is a parameter which varies from 0 to *norbit*.

```
In[78]:= n[z_] := IntegerPart[z]
```

```
In[79]:= zf[z_] := FractionalPart[z]
```

```
In[80]:= ua[z_] := u1 (1 - 2 zf[z]) + u2 2 zf[z]
```

```
In[81]:= ub[z_] := u1 (2 zf[z] - 1) + 2 u2 (1 - zf[z])
```

```
In[82]:= u[z_] := If[zf[z] < .5, ua[z], ub[z]]
```

```
In[83]:= phia[z_] := 2 (n[z]) delphi + theta[u[z], ε, ℓ, u1]
```

```
In[84]:= phib[z_] := 2 (n[z] + 1) delphi - theta[u[z], ε, ℓ, u1]
```

```
In[85]:= accphi[z_] := If[zf[z] < .5, phia[z], phib[z]]
```

```
In[86]:= x[z_] := Cos[accphi[z]] / u[z]
```

```
In[87]:= y[z_] := Sin[accphi[z]] / u[z]
```

■ Displaying the Results:

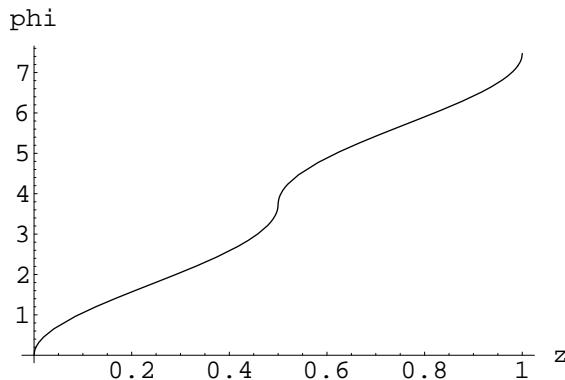
accphi is the accumulated angle as a function of position along the orbit, as measured by the parameter *z* defined above to vary from 0 to *norbit*, so that *z*=1 is the end of the first orbit, *z*=2 is the end of the second, etc.

```
In[88]:= If[testa, norbit = norbit, norbit = 1]
```

```
Out[88]= 3
```

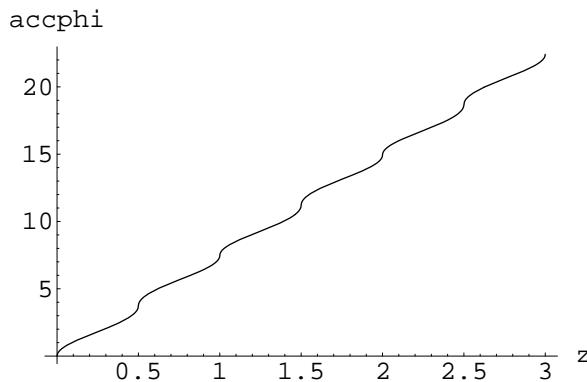
```
In[89]:= If[testd, norbit = .5]
```

```
In[90]:= If[norbit > 1, Plot[accphi[z], {z, 0, 1}, AxesLabel -> {"z", "phi"}]]
```



```
Out[90]= - Graphics -
```

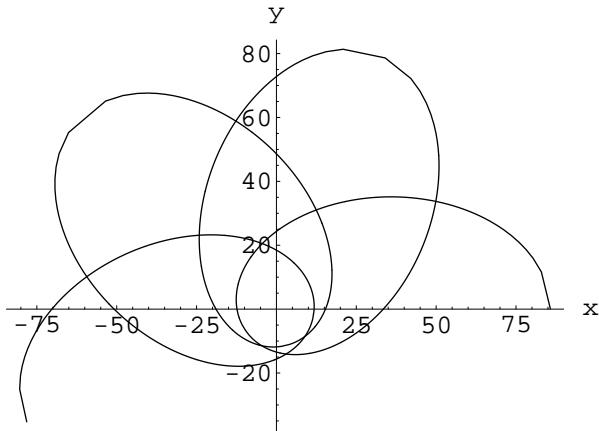
```
In[91]:= Plot[accphi[z], {z, 0, norbit}, AxesLabel -> {"z", "accphi"}]
```



```
Out[91]= - Graphics -
```

Calculating and plotting the orbit: Two parameters specify how accurately the orbit is calculated: **PlotDivision** specifies the number of plot points, **MaxBend** specifies the maximum bending angle permitted. For shorter calculations and rougher orbits decrease **PlotDivision** and increase **MaxBend**. For longer calculations and more accurate orbits increase **PlotDivision** and decrease **MaxBend**.

```
In[92]:= graph = ParametricPlot[{x[t], y[t]}, {t, 0, norbit}, MaxBend -> .1,
    PlotDivision -> 50, AspectRatio -> Automatic, AxesLabel -> {"x", "y"}]
```



```
Out[92]= - Graphics -
```

If you want to output a copy of the orbit as an eps file called orbit.eps you can include the statement:
`Display["orbit.eps", graph, "EPS"].`