

String Theory 230A Homework # 1 Solutions

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Problem 2 - Geodesic Equations

Consider a relativistic point particle in $d + 1$ dimensions, with action

$$S_\tau = -m \int \sqrt{-g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau}} d\tau = -m \int ds$$

(a)

Under the reparametrization

$$\tau' = \tau'(\tau)$$

the action becomes

$$\begin{aligned} S_\tau &= -m \int \sqrt{-g_{\mu\nu} \frac{dx^\mu}{d\tau'} \frac{dx^\nu}{d\tau'}} \left(\frac{d\tau'}{d\tau} \right) d\tau \\ &= -m \int \sqrt{-g_{\mu\nu} \frac{dx^\mu}{d\tau'} \frac{dx^\nu}{d\tau'}} d\tau' \\ &= S_{\tau'} \end{aligned}$$

where we have repeatedly used the chain rule. So the action is invariant under reparametrization.

(b)

If we choose

$$\frac{ds}{d\tau} = \left(-g_{\mu\nu} \frac{dx^\mu}{d\tau} \frac{dx^\nu}{d\tau} \right)^{1/2}$$

then

$$\left(-g_{\mu\nu} \frac{dx^\mu}{ds} \frac{dx^\nu}{ds} \right) = 1$$

by part (a).

(c)

We compute the following variations

$$\begin{aligned} \frac{\partial S}{\partial \dot{x}^\alpha} &= g_{\alpha\mu} \dot{x}^\mu \\ \frac{\partial S}{\partial x^\alpha} &= \frac{1}{2} (\partial_\alpha g_{\mu\nu}) \dot{x}^\mu \dot{x}^\nu \\ \frac{d}{ds} \left(\frac{\partial S}{\partial \dot{x}^\alpha} \right) &= (\partial_\nu g_{\alpha\mu}) \dot{x}^\mu \dot{x}^\nu + g_{\alpha\mu} \ddot{x}^\mu \end{aligned}$$

where we have set

$$\left(-g_{\mu\nu} \frac{dx^\mu}{ds} \frac{dx^\nu}{ds}\right) = 1$$

after computing the variation. After symmetrizing

$$(\partial_\nu g_{\alpha\mu}) \dot{x}^\mu \dot{x}^\nu = \frac{1}{2} (\partial_\nu g_{\alpha\mu} + \partial_\mu g_{\alpha\nu}) \dot{x}^\mu \dot{x}^\nu$$

the Euler-Lagrange equation yields

$$\ddot{x}_\mu + \frac{1}{2} (g_{\alpha\mu,\nu} + g_{\alpha\nu,\mu} - g_{\mu\nu,\alpha}) \dot{x}^\mu \dot{x}^\nu = 0.$$

Finally we raise the index on \ddot{x}_μ by contracting both sides with $g^{\alpha\beta}$ yielding the geodesic equation

$$\boxed{\ddot{x}^\beta + \Gamma_{\mu\nu}^\beta \dot{x}^\mu \dot{x}^\nu = 0.}$$

String Theory 229A Homework Solution Set #1

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1. (15 Points)

(a) The Lagrangian for the point particle in an EM field is, from the lecture notes:

$$\mathcal{L} = -m\sqrt{-\dot{x}^\mu \dot{x}_\mu} \pm eA_\mu \dot{x}^\mu \quad (1)$$

Therefore the Euler-Lagrange equations give the equations of motion:

$$\frac{\partial}{\partial \tau} \frac{\partial \mathcal{L}}{\partial \dot{x}^\mu} = \frac{\partial \mathcal{L}}{\partial x^\mu} \quad (2)$$

$$\frac{\partial}{\partial \tau} \left(\frac{m\dot{x}_\mu}{\sqrt{-\dot{x}^\nu \dot{x}_\nu}} \pm eA_\mu \right) = \pm e \frac{\partial A_\nu}{\partial x^\mu} \dot{x}^\nu \quad (3)$$

$$m\dot{x}_\mu \pm e \frac{\partial x^\nu}{\partial \tau} \frac{\partial}{\partial x^\nu} A_\mu = \pm e \frac{\partial A_\nu}{\partial x^\mu} \dot{x}^\nu \quad (4)$$

$$(5)$$

$$m\dot{x}_\mu \pm e(\partial_\nu A_\mu - \partial_\mu A_\nu) \dot{x}^\nu = 0 \quad (6)$$

$$m\dot{x}_\mu = \pm e F_{\mu\nu} \dot{x}^\nu \quad (7)$$

(b) Instead of the hint in the problem, we can obtain the identical system by choosing a particular parameterization, namely we choose τ such that $-\dot{x}^\mu \dot{x}_\mu = m^2$. Taking this constraint, we see that the equation of motion derived in part a becomes simply:

$$\ddot{x}_\mu = \pm e F_{\mu\nu} \dot{x}^\nu \quad (8)$$

$$m^2 = -\dot{x}^\mu \dot{x}_\mu \quad (9)$$

Choose the positive sign for e so we stop carrying around those pesky \pm 's. Plugging in for the explicit form of $F_{\mu\nu}$, we get

$$\ddot{x}_1 = \ddot{x}_2 = 0 \quad (10)$$

$$\ddot{x}_3 = -eB\dot{x}_4 \quad (11)$$

$$\ddot{x}_4 = eB\dot{x}_3 \quad (12)$$

plus the constraint above. These are a standard set of coupled differential equations, their solutions are well know,

and easily derived:

$$x_1 = a_1 + a_2\tau \quad (13)$$

$$x_2 = b_1 + b_2\tau \quad (14)$$

$$x_3 = c_1 + \frac{c_2}{eB} \sin(eB\tau) + \frac{c_3}{e^2B^2} \cos(eB\tau) \quad (15)$$

$$x_4 = c_4 - \frac{c_2}{eB} \cos(eB\tau) + \frac{c_3}{e^2B^2} \sin(eB\tau) \quad (16)$$

$$(17)$$

If we assume simplifying initial conditions such that $a_1 = 0$ and $a_2 = 1$, then $\tau = x_1 = t$, and we can rewrite the solution in a more standard classical form:

$$\vec{x}(t) = (b_1 + b_2t)\hat{x}_2 + \left(c_1 + \frac{c_2}{eB} \sin(eBt) + \frac{c_3}{e^2B^2} \cos(eBt) \right) \hat{x}_3 \quad (18)$$

$$+ \left(c_4 - \frac{c_2}{eB} \cos(eBt) + \frac{c_3}{e^2B^2} \sin(eBt) \right) \hat{x}_4 \quad (19)$$

This motion is now immediately identified as a helix in the 3-4 plane, as we would expect for the motion of a charged particle in a constant magnetic field.

(c) Starting from the covariant form of the lagrangian:

$$\mathcal{L} = -\frac{1}{2}(\dot{x}^\mu \dot{x}_\mu - m^2) \pm eA_\mu \dot{x}^\mu \quad (20)$$

(note there is a typo in the lecture notes with the sign of m^2) The conjugate momenta of the x^μ is trivially found:

$$p_\mu \equiv \frac{\partial \mathcal{L}}{\partial \dot{x}^\mu} = -\dot{x}_\mu \pm eA_\mu \quad (21)$$

So the Hamiltonian is:

$$H \equiv p^\mu \dot{x}_\mu - \mathcal{L} \quad (22)$$

$$= -p^\mu (p_\mu \mp eA_\mu) + \frac{1}{2}(\dot{x}^\mu \dot{x}_\mu - m^2) \mp eA^\mu \dot{x}_\mu \quad (23)$$

$$= -p^\mu (p_\mu \mp eA_\mu) + \frac{1}{2}((p_\mu \mp eA_\mu)(p^\mu \mp eA^\mu) - m^2) + eA^\mu (p_\mu \mp eA_\mu) \quad (24)$$

$$= -\frac{1}{2}(p_\mu \mp eA_\mu)(p^\mu \mp eA^\mu) - \frac{1}{2}m^2 \quad (25)$$

Because the original action was manifestly covariant, we know that the constraint will just be $H = 0$ or $H|\psi\rangle = 0$ on quantum states. Let's verify this. We know that the constraint with $\eta = 1$ from the η equations of motion imply that $m^2 = -\dot{x}^\mu \dot{x}_\mu$. Therefore, in terms of the momentum variables, $m^2 + (p_\mu \mp eA_\mu)(p^\mu \mp eA^\mu) = 0$, which is exactly $H = 0$ as we derived above.

So, defining $D_\mu = \partial_\mu \mp ieA_\mu$, if we replace the momentum in the normal way $p_\mu \rightarrow -i\partial_\mu$, the quantum constraint is:

$$H|\psi\rangle = 0 \quad (26)$$

$$(-i\partial_\mu \mp eA_\mu)(-i\partial^\mu \mp eA^\mu) + m^2|\psi\rangle = 0 \quad (27)$$

$$(\partial_\mu \mp ieA_\mu)(\partial^\mu \mp ieA^\mu) - m^2|\psi\rangle = 0 \quad (28)$$

$$(D_\mu D^\mu - m^2)|\psi\rangle = 0 \quad (29)$$

That is the Klein-Gordon equation.

2. 15 Points

(a) The action of the point particle coupled to a background of the form given in the problem becomes:

$$S = -m \int d\tau \sqrt{\dot{x}^+ \dot{x}^- + \beta x_\perp^2 \dot{x}^{+2} - \dot{x}_\perp^2} \quad (30)$$

Using reparameterization invariance to choose $x^+ = \tau$, it becomes:

$$S = -m \int d\tau \sqrt{\dot{x}^- + \beta x_\perp^2 - \dot{x}_\perp^2} \quad (31)$$

Now we follow the lecture notes, and find the Hamiltonian corresponding to this system. First we get the conjugate momentum:

$$p_\perp = \frac{\partial \mathcal{L}}{\partial \dot{x}_\perp} = \frac{m \dot{x}_\perp}{\sqrt{A}} \quad (32)$$

$$p_- = \frac{\partial \mathcal{L}}{\partial \dot{x}^-} = -\frac{m}{2\sqrt{A}} \quad (33)$$

$$A = \dot{x}^- + \beta x_\perp^2 - \dot{x}_\perp^2 \quad (34)$$

Note that

$$\sqrt{A} = -\frac{m}{2p_-} \quad (35)$$

$$\dot{x}_\perp = -\frac{p_\perp}{2p_-} \quad (36)$$

$$\dot{x}^- = \frac{m^2 + p_\perp^2}{4p_-^2} - \beta x_\perp^2 \quad (37)$$

The Hamiltonian is then

$$H = p_- \dot{x}^- + p_\perp \dot{x}_\perp - \mathcal{L} \quad (38)$$

$$= \frac{m^2 + p_\perp^2}{4p_-} - p_- \beta x_\perp^2 - \frac{p_\perp^2}{2p_-} - \frac{m^2}{2p_-} \quad (39)$$

$$= -\frac{1}{4} \left(\frac{m^2 + p_\perp^2}{p_-} \right) - p_- \beta x_\perp^2 \quad (40)$$

We know that p_- is constant since it is related to p^+ by $p_- = -\frac{1}{2}p^+$, so in terms of p^+ we simply have

$$H = \frac{m^2}{2p^+} + \frac{p_\perp^2}{2p^+} + \frac{\beta p^+ x_\perp^2}{2} \quad (41)$$

We recognize this as simply the hamiltonian of the simple harmonic oscillator, with β being the frequency, and p^+ playing the role of mass.

(b) The equations of motion (given the hamiltonian above) are trivially:

$$\ddot{x}_\perp + \beta x_\perp = 0 \quad (42)$$

This is simply the standard harmonic oscillator solution for the directions perpendicular to the light cone.

$$x_\perp(\tau) = c_1 \cos(\sqrt{\beta}\tau) + c_2 \sin(\sqrt{\beta}\tau) \quad (43)$$

For the one extra direction x^- , we use the fact that p_- is a constant, therefore A is a constant as defined above, call it d :

$$A = d \quad (44)$$

$$\dot{x}^- = d + \dot{x}_\perp^2 - \beta x_\perp^2 \quad (45)$$

$$\dot{x}^- = d + (c_2^2 - c_1^2) \cos(2\sqrt{\beta}\tau) - 2c_1 c_2 \sin(2\sqrt{\beta}\tau) \quad (46)$$

$$x^- = e + d\tau + \frac{(c_2^2 - c_1^2)}{2\sqrt{\beta}} \sin(2\sqrt{\beta}\tau) + \frac{c_1 c_2}{\sqrt{\beta}} \cos(2\sqrt{\beta}\tau) \quad (47)$$

with a little bit of algebra, where c_1, c_2, d , and e are constants.

(c) With the Lagrange multiplier, the action becomes:

$$S = \frac{1}{2} \int d\tau (\eta^{-1}(\dot{x}^- + \beta x_\perp^2 - \dot{x}_\perp^2) + \eta m^2) \quad (48)$$

The η equation of motion is:

$$\eta^2 = \frac{\dot{x}^- + \beta x_\perp^2 - \dot{x}_\perp^2}{m^2} \quad (49)$$

so upon substitution back into the action we recover the original action.

The Euler Lagrange equation of motion for x_\perp is:

$$\eta^{-1} \ddot{x}_\perp + \eta^{-1} \beta x_\perp = 0 \quad (50)$$

(d) Choosing $\eta = 1$, the equation of motion becomes:

$$\ddot{x}_\perp + \beta x_\perp = 0 \quad (51)$$

exactly as we had before (actually, we could have chosen η as anything except 0 and this equation of motion would have been the same, but the next equation would not be) Combined with the constraint

$$m^2 = \dot{x}^- + \beta x_\perp^2 - \dot{x}_\perp^2 \quad (52)$$

we have the exact same set of equations as we did in part b, and therefore the solutions will be identical as well.

(e) Let's see the hamiltonian again:

$$H = \frac{m^2}{2p^+} + \frac{p_\perp^2}{2p^+} + \frac{\beta p^+ x_\perp^2}{2} \quad (53)$$

As stated above, this is nothing but the harmonic oscillator hamiltonian in the $D - 2$ directions perpendicular to the light cone. So upon quantization, you will get $D - 2$ uncoupled harmonic oscillators, each contributing energy $\sqrt{\beta}(n + \frac{1}{2})$ for some integer quantum number n . Since you have $D - 2$ oscillators, you have $D - 2$ quantum numbers to describe the state.